

# Thermodynamics of holographic models for QCD in the Veneziano limit

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[TA, Järvinen, Kajantie, Kiritsis, Tuominen arXiv:1210.4516]

# Overview

- Veneziano QCD
- The model and determining the potentials
- Computing thermodynamics
- Results
- Outlook

# Veneziano QCD

Veneziano QCD is a YM theory with  $N_c$  colors and  $N_f$  fermion flavors, at the limit  $N_c, N_f \rightarrow \infty$  but  $x_f \equiv \frac{N_f}{N_c}$  constant.

## Veneziano QCD

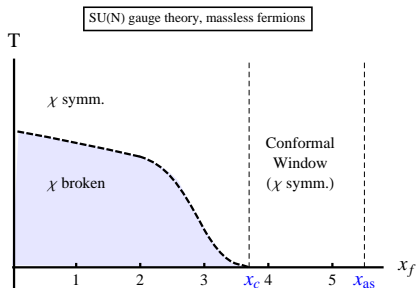
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The holographic model and its vacuum structure described in arXiv:1112.1261, and in the previous talk by Järvinen. For studying the thermodynamics, we add a black hole to the bulk. The model stays the same, but the metric ansatz now becomes

$$ds^2 = b^2(r) \left[ -f(r)dt^2 + d\mathbf{x}^2 + \frac{dr^2}{f(r)} \right], \quad (1)$$

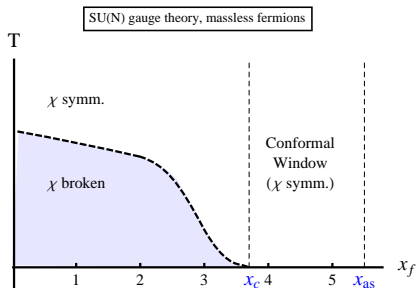
# What to expect

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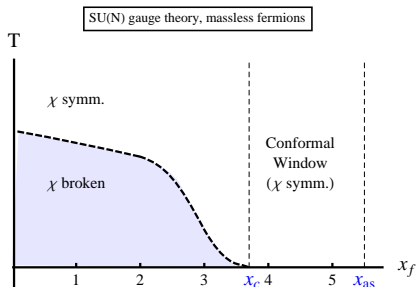
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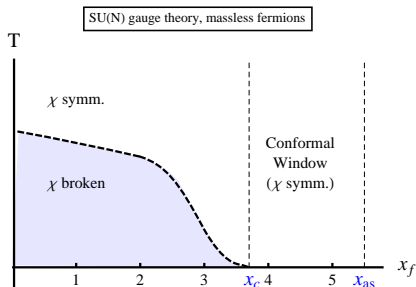
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What we expect from QCD -like thermodynamics:



- Conformal window between  $x_c \approx 4$  and  $x_{as} = 11/2$
- A deconfinement/chiral symmetry restoring transition at  $x_f < x_c$
- Miransky scaling at  $x_f \lesssim x_c$



## Action

To recap the setup, the gravity action is

$$S = M^3 N_c^2 \int d^5 x \mathcal{L} \equiv \frac{1}{16\pi G_5} \int d^5 x \mathcal{L}, \quad (2)$$

where

$$\mathcal{L} = \left[ \sqrt{-g} \left( R - \frac{4}{3} \frac{(\partial\lambda)^2}{\lambda^2} + V_g(\lambda) \right) - V_f(\lambda, \tau) \sqrt{\det(g_{ab} + \kappa(\lambda, \tau) \partial_a \tau \partial_b \tau)} \right]. \quad (3)$$

The metric Ansatz is

$$ds^2 = b^2(r) \left[ -f(r) dt^2 + d\mathbf{x}^2 + \frac{dr^2}{f(r)} \right], \quad (4)$$

and the two scalar functions,  $1/\lambda$  sourcing  $F^2$  and  $\tau$  sourcing  $\langle \bar{q}q \rangle$ , are

$$\lambda = \lambda(r) = e^{\phi(r)}, \quad \tau = \tau(r). \quad (5)$$

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- IR limit: confinement and  $\tau$  divergence



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- The string frame to Einstein frame conversion factor  $\kappa$  may be corrected by a logarithmic factor.

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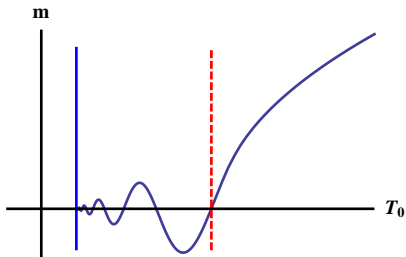
- chiral symmetry  $\Rightarrow \tau(r) \equiv 0$
- chiral symmetry broken  $\Rightarrow \tau(r) \neq 0$ , but  $m_q = 0$  determines  $\tau_h$  as a function of  $\lambda_h$

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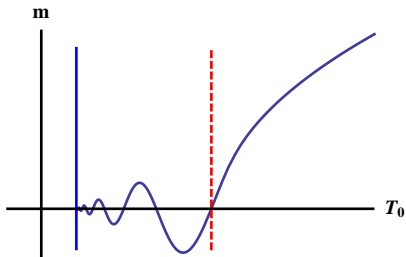
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- However, only the one with largest  $\tau_h$  is thermodynamically relevant (smallest free energy)
- Overall, two separate branches of black hole solutions, one with  $\tau \equiv 0$ , and one with a dynamic tachyon.

# Thermodynamics

$$T = -\frac{1}{4\pi} f'(r_h(\lambda_h))$$

$$s = \frac{1}{4G_5} b^3(\lambda_h)$$

$$p = -F = \frac{1}{4G_5} \int_{\lambda_h}^{\lambda_*} d\lambda_h \left( -\frac{dT}{d\lambda_h} \right) b^3(\lambda_h) + p_0$$

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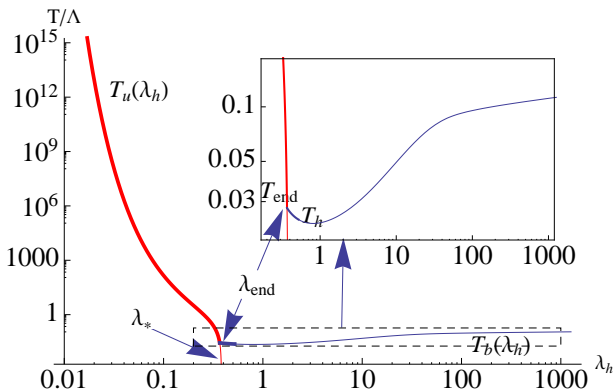
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Note that  $c_s^2 > 0$  requires that  $\frac{dT}{d\lambda_h} < 0$ .

# Temperature

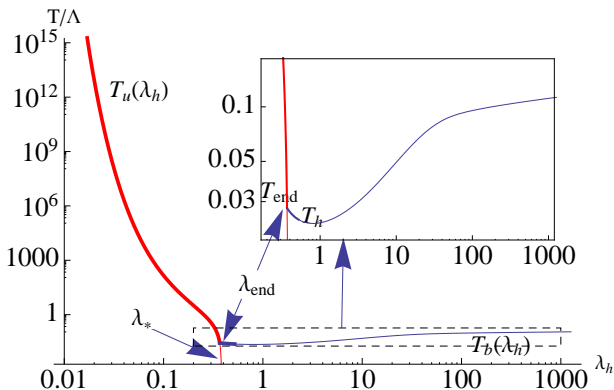
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Below  $\lambda_h < \lambda_{\text{end}}$  the  $\tau_h \neq 0$  phase doesn't exist  $\Rightarrow$  no chiral symmetry breaking at high  $T$

## Computing pressure

To compute the pressure, we need to fix the integration constants in the two branches:

- At the  $\lambda_h \rightarrow \infty$  limit, the  $\tau \neq 0$  solution becomes the  $\tau \neq 0$  vacuum solution, i.e.  $p_b(\lambda_h = \infty) = 0$ .

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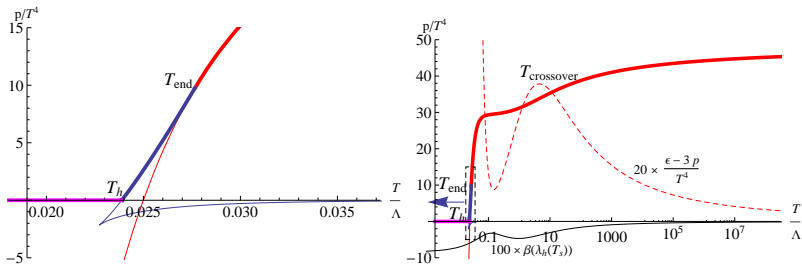
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- Require that the difference in free energy between these is the same as between the corresponding vacuum solutions  
 $\Rightarrow$  numerically equivalent to requiring that  $p_u(\lambda_{\text{end}}) = p_b(\lambda_{\text{end}})$

# Pressure

Typical  $p(T)$ , with  $x_f = N_f/N_c = 3$ :



Phase transitions in order of decreasing  $T$ :

- A crossover around  $T_{\text{crossover}}$
- The second order chiral symmetry breaking transition at  $T_{\text{end}}$
- The confining, or hadronization, transition at  $T_h$  from the black hole phase to the hadron gas phase

## Numerics

From the above, calculate the  $(x_f, T)$  phase diagram and other thermodynamic variables for a comprehensive subset of all choices of potential. A numerical code which near-automatically gives the full thermodynamics given the potentials as inputs allowed us to explore ten different potentials with reasonable time and effort.

## Numerics

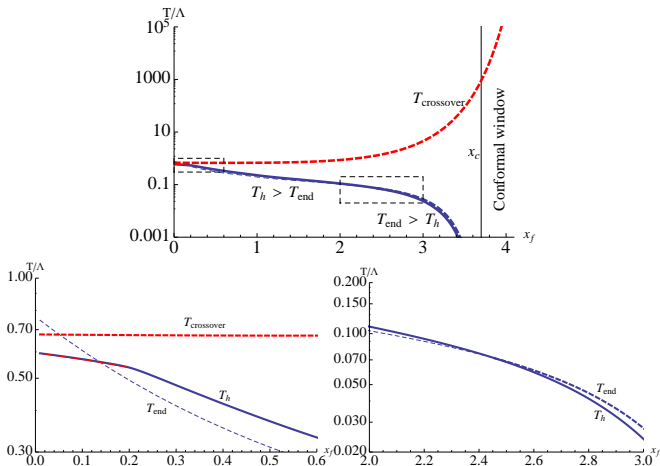
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Some examples to follow. Legend:

- blue lines: transitions which involve either the  $\tau \neq 0$  -branch or the  $\tau \neq 0$  vacuum solution
- red lines: transitions which only involve the  $\tau = 0$  -solution.
- thick lines: transitions between two stable phases
- thin lines: transitions between two metastable phases

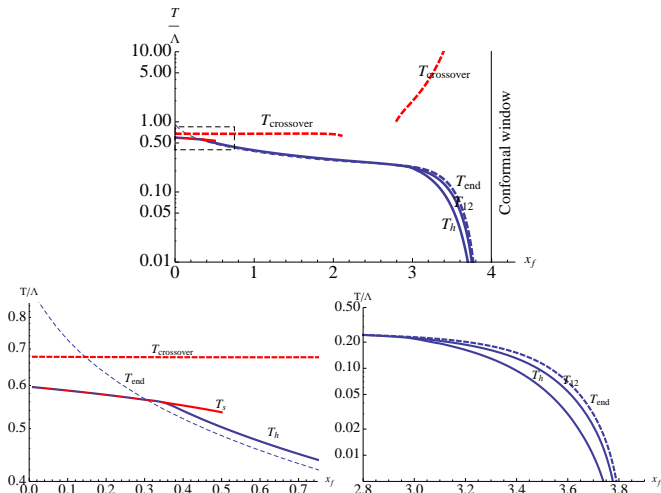


## Potential II, $W_0$ SB-normalized



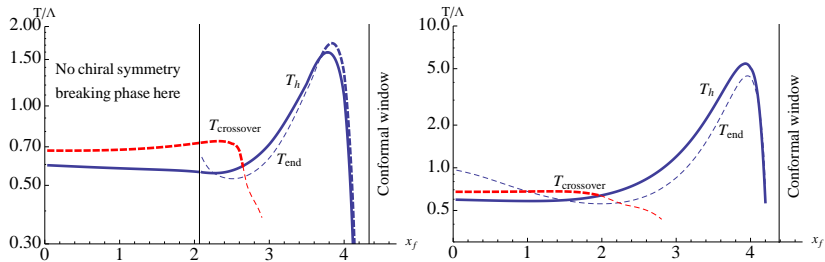
Reasonable phase diagram, but meson masses go as  $M_n \sim n$ , so not a good QCD model.

# Potential I, $W_0 = 12/11$

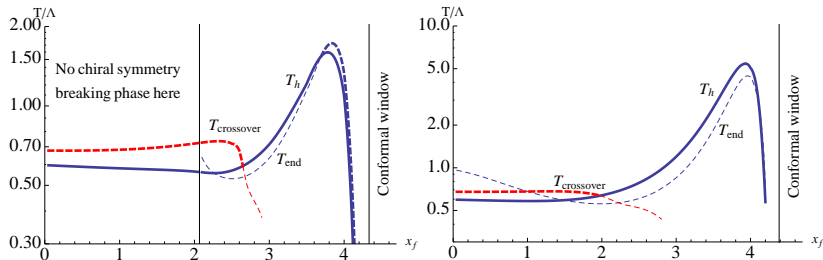


The most complicated phase diagram we found.

# Potentials $I_*$ and $II_*$

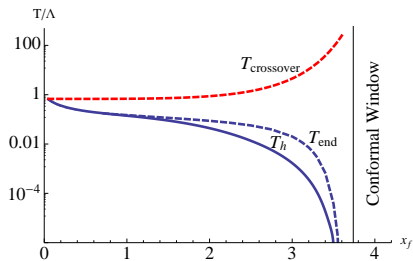


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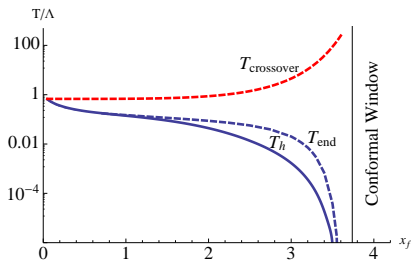


$I_*$  has no  $\chi_S B$  at low- $x_f$ ,  $II_*$  has the same spectrum problem as  $II$ .

# Potential I with log-modified $\kappa$

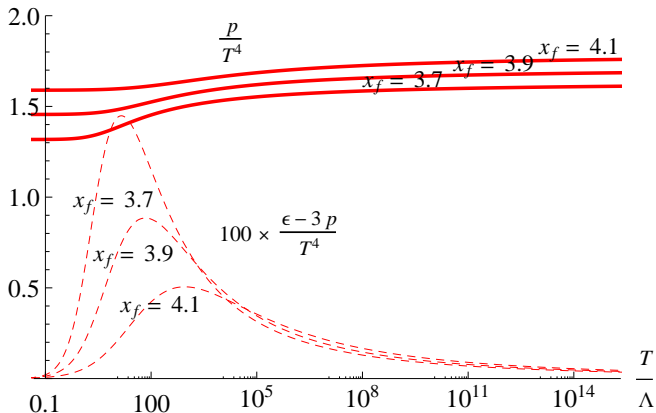


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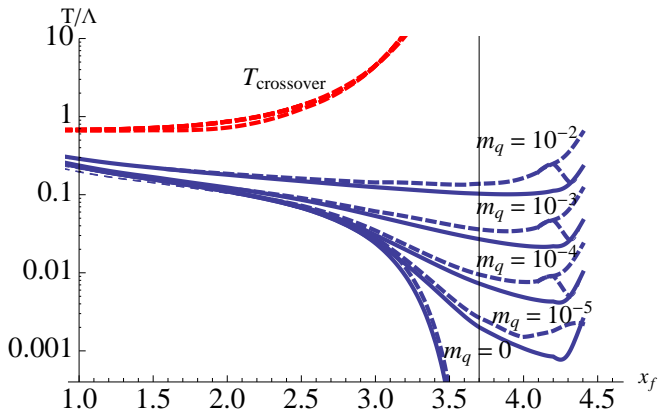
- phase diagram compatible with expectations for QCD-like theory
- meson trajectory is Regge-like
- $\Rightarrow$  best candidate for modeling QCD-like theories (for now, Kiritsis et. al. working further on this)

## The conformal window



No  $\tau \neq 0$  solution in the conformal window  $\Rightarrow$  thermodynamics in the conformal window is qualitatively independent of the choice of potential.

# Finite mass (potential II, $W_0$ SB -normalized)





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- A region with a walking coupling constant, i.e. quasi-conformality, when  $x_f$  is slightly below  $x_c$
- at small  $x_f$ , chiral symmetry restoration always coincides with deconfinement.

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That's all, folks! Thank you!