

# Cosmology and the Dark Energy Survey

400

Adam Amara ETH Zurich Dark Energy Survey & Latest Results

Cosmic Concordance?

Precision Measurements Looking Forward

# The Dark Energy Survey

#### DES First Light 12 September 2012







# Footprint



year 0 (science validation) - 180 deg<sup>2</sup> (full depth) year 1 - 2500 deg<sup>2</sup> (half depth) 5 years - 5000 deg<sup>2</sup>

# History of Light Bending

Dyson, Eddington & Davidson 1919

1915: General Relativity
1919: Eclipse Experiment
1937: Galaxies as Lens (*Zwicky*)
1979: First Galaxy Lens





## **Examples of Gravitational Lenses**



#### Bridle, AA + 2008 AA 2011

## **Galaxy Shapes**



- Key is a well-behaved PSF
- Precisions hardware
- Analysis methods

Voigt, Bridle, AA+2012 Kacprzak,..., AA+2012 Refregier, Kacprzak, AA+ 2012 Refregier & AA 2014 AA+ 2010 Cypriano, AA+ 2010 Paulin-Henriksson, Refregier & AA 2009 Paulin-Henriksson, AA+ 2008

## Photometric Redshifts



.8 AA & Refregier 2007 Abdalla, AA+ 2008 Bordoloi, AA & Lilly 2010 Bordoloi, AA & Lilly 2012 Bonnett, Troxel, Hartley, AA+ 2015

### Weak Lensing Correlations and Mass Reconstruction



Fu et al 2007 - CFHTLS

Massey et al 2007 - COSMOS

# **DES: Dark Matter Mapping**



Chang,..., AA+ PRL (2015) Vikram, Chang..., AA+ PRD (2015) Chang, Pujol, Gaztañaga, AA+ (2016) Pujol, Chang, Gaztañaga, AA+ (2016) 22

# Science Verification Results

Cosmlogy	DES Collaboration (arXiv:1507.05603)
Shear Catalogs	Jarvis et al (arXiv:1507.05603)
Photometric redshift	Bonnett et al (arXiv:1507.05909)
Systematics maps	Leistedt et al (arXiv:1507.05647)
Shear Power Spectra	Becker et al (arXiv:1507.05598)

## **DES: Dark Matter Statistics**



# Robustness



# **Cosmology Parameters**



#### The Dark Energy Survey Collaboration (2015)

Dark Energy Survey & Latest Results

Cosmic Concordance?

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# Cosmological Probes



## **Cosmic Microwave Background**



	Parameter	WMAP		+eCMB +eCM		AB+BAO +eCMB+A		H <sub>0</sub> -	+eCMB+BAO+H <sub>0</sub>			
	Fit parameters											
	$\Omega_b h^2$	0.0226	$54 \pm 0.00050$	$0.02229 \pm 0$	.00037	0.02211	$\pm 0.00034$	$0.02244 \pm 0.0$	)0035 (	$0.02223 \pm 0.00033$		
	$\Omega_c h^2$	0.113	$38 \pm 0.0045$	$0.1126 \pm 0$	.0035	0.1162	$2 \pm 0.0020$	$0.1106 \pm 0.0$	0030	$0.1153 \pm 0.0019$		
	$\Omega_{\Lambda}$	0.72	$21 \pm 0.025$	$0.728 \pm 0$	.019	0.707	$2 \pm 0.010$	$0.740 \pm 0.0$	)15	$0.7135_{-0.0096}^{+0.0095}$		
	$10^9\Delta_R^2$	2.4	$11 \pm 0.10$	$2.430 \pm 0$	.084	2.48	$84^{+0.073}_{-0.072}$	$2.396_{-0.0}^{+0.01}$	78	$2.464 \pm 0.072$		
Q	$n_s$	0.97	$72 \pm 0.013$	$0.9646 \pm 0$	.0098	0.957	$79^{+0.0081}_{-0.0082}$	$0.9690\substack{+0.00\\-0.0}$	)91 090	$0.9608 \pm 0.0080$		
J	τ	0.08	$39 \pm 0.014$	$0.084 \pm 0$	.013	0.07	$79^{+0.011}_{-0.012}$	$0.087 \pm 0.0$	013	$0.081 \pm 0.012$		
	Derived parameters											
	$t_0$ (Gyr)	$13.74\pm0.11$		$13.742 \pm 0$	.077	13.800	$0 \pm 0.061$	$13.702 \pm 0.0$	)69	$13.772 \pm 0.059$		
	$H_0 ({\rm km}{\rm s}^{-1}{\rm Mpc}^{-1})$	70	$.0 \pm 2.2$	$70.5 \pm 1$	.6	68.76	$5 \pm 0.84$	$71.6 \pm 1.4$	4	$69.32 \pm 0.80$		
	$\sigma_8$	0.82	$21 \pm 0.023$	$0.810 \pm 0$	.017	0.82	$22^{+0.013}_{-0.014}$	$0.803 \pm 0.0$	)16	$0.820^{+0.013}_{-0.014}$		
	$\Omega_b$	$0.0463 \pm 0.0024$		$0.0449 \pm 0$	.0018	0.04678	$3 \pm 0.00098$	$0.0438 \pm 0.0$	0015 (	$0.04628 \pm 0.00093$		
	$\Omega_c$	$0.233 \pm 0.023$		$0.227 \pm 0.017$		$0.2460 \pm 0.0094$		$0.216 \pm 0.0$	014	$0.2402^{+0.0088}_{-0.0087}$		
	Zeq	$3265^{+106}_{-105}$		$3230 \pm 81$		$3312\pm48$		$3184 \pm 70$	I	$3293 \pm 47$		
	Zreion	10	$.6 \pm 1.1$	$10.3 \pm 1$	.1	10.0	$0 \pm 1.0$	$10.5 \pm 1.1$	1	$10.1 \pm 1.0$		
				Planck+WP+highI Planck+le		ensing+WP+highI Planck		WP+highL+BAO				
		1		1 iun		<u>, , , , , , , , , , , , , , , , , , , </u>						
	Parameter	Best fit	68% limits	Best fit	68% li	mits	Best fit	68% limits	Best fit	68% limits		
	$\Omega_{ m b}h^2$	0.022032	$0.02205 \pm 0.00028$	0.022069	$0.02207 \pm$	0.00027	0.022199	$0.02218 \pm 0.00026$	0.022161	$0.02214 \pm 0.00024$		
	$\Omega_{\rm c}h^2$	0.12038	$0.1199 \pm 0.0027$	0.12025	$0.1198 \pm$	0.0026	0.11847	$0.1186 \pm 0.0022$	0.11889	$0.1187 \pm 0.0017$		
	$100\theta_{\rm MC}$	1.04119	$1.04131 \pm 0.00063$	1.04130	$1.04132 \pm$	0.00063	1.04146	$1.04144 \pm 0.00061$	1.04148	$1.04147 \pm 0.00056$		
	τ	0.0925	$0.089^{+0.012}_{-0.014}$	0.0927	0.091_	0.013 0.014	0.0943	$0.090^{+0.013}_{-0.014}$	0.0952	$0.092 \pm 0.013$		
<	$n_{\rm s}$	0.9619	$0.9603 \pm 0.0073$	0.9582	$0.9585 \pm$	0.0070	0.9624	$0.9614 \pm 0.0063$	0.9611	$0.9608 \pm 0.0054$		
`	$\ln(10^{10}A_{\rm s})$	3.0980	$3.089^{+0.024}_{-0.027}$	3.0959	$3.090 \pm$	0.025	3.0947	$3.087 \pm 0.024$	3.0973	$3.091 \pm 0.025$		
	$\overline{\Omega_{\Lambda}$	0.6817	$0.685^{+0.018}_{-0.016}$	0.6830	0.685_	0.017 0.016	0.6939	$0.693 \pm 0.013$	0.6914	$0.692 \pm 0.010$		
	$\sigma_8$	0.8347	$0.829 \pm 0.012$	0.8322	$0.828 \pm$	0.012	0.8271	$0.8233 \pm 0.0097$	0.8288	$0.826 \pm 0.012$		
	$z_{re}$	11.37	$11.1 \pm 1.1$	11.38	11.1 ±	1.1	11.42	$11.1 \pm 1.1$	11.52	$11.3 \pm 1.1$		
	$H_0$	67.04	$67.3 \pm 1.2$	67.15	67.3 ±	1.2	67.94	$67.9 \pm 1.0$	67.77	$67.80 \pm 0.77$		
	Age/Gyr	13.8242	$13.817\pm0.048$	13.8170	13.813 ±	0.047	13.7914	$13.794 \pm 0.044$	13.7965	$13.798 \pm 0.037$		
	$100\theta_*$	1.04136	$1.04147 \pm 0.00062$	1.04146	$1.04148 \pm$	0.00062	1.04161	$1.04159 \pm 0.00060$	1.04163	$1.04162 \pm 0.00056$		
	<i>r</i> <sub>drag</sub>	147.36	$147.49\pm0.59$	147.35	147.47 =	± 0.59	147.68	$147.67\pm0.50$	147.611	$147.68\pm0.45$		

### WMAP 9

Hinshaw+ (2013)

Planck (2013)



What are the central values?

Have things changed?

Are things consistent?



### RELATIVE ENTROPY

### Constraints from Data A neter transformations



Kullback & Leibler (1951)





### NORMAL DISTRIBUTIONS & LINEAR MODEL

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	Fit parameters											
	$\Omega_b h^2$	0.0226	$54 \pm 0.00050$	$0.02229 \pm 0$	.00037	0.02211	$\pm 0.00034$	$0.02244 \pm 0.000000000000000000000000000000000$	)0035 (	$0.02223 \pm 0.00033$		
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### WMAP 9

Hinshaw+ (2013)

Planck (2013)

	D	$\langle D \rangle$	S	$S/\sigma(D)$	p-value <sup>d</sup>
BOOMERANG $\rightarrow$ WMAP 9	22.5	18.4	4.1	1.6	0.07
WMAP 9 $\rightarrow$ WMAP 9 + SPT	4.3	2.1	2.2	2.1	0.04
WMAP 9 $\rightarrow$ Planck + WP	29.8	7.9	21.9	6.5	0.0002

Unit of bits

BOOMERANG: MacTavish et al. (2003) WMAP 3, 5, 7, 9: Spergel et al. (2007), Dunkley et al. (2009), Larson et al. (2011), and Bennett et al. (2013) WP: WMAP 9 polarisation data SPT: Story et al. (2013) Planck: Ade et al. (2013)



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What about the other probes?

#### Grandis et al (2015)



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# Current and Planned Experiments

Survey	Start (rough dates)
COSMOS*	2003
CFHTLS	2003
Pan-STARRS1	2009
KIDS	2011
DES*	2012
HALO (balloon)*	>2020
LSST*	>2020
Euclid*	>2020
WFIRST	>2020

![](_page_38_Picture_2.jpeg)

![](_page_38_Picture_3.jpeg)

![](_page_38_Picture_4.jpeg)

#### Bridle, Amara+ 2008 Amara 2011

## **Galaxy Shapes**

![](_page_39_Figure_2.jpeg)

Analysis methods

Voigt, Bridle, Amara+2012 Kacprzak,..., Amara+2012 Refregier, Kacprzak, Amara+2012 Refregier & Amara 2014

Amara+ 2010 Cypriano, Amara+ 2010 Paulin-Henriksson, Refregier & Amara 2009 Paulin-Henriksson, Amara+ 2008

### Toy Model: Measuring the Size of a 2D Gaussian

![](_page_40_Figure_1.jpeg)

![](_page_40_Picture_2.jpeg)

Refregier, AA, + 2013

# Measurement Biases

![](_page_41_Figure_1.jpeg)

$$\delta a_i \simeq -\frac{1}{2} F_{ij} F_{kl} B_{jkl} \propto 1/\text{SNR}^2$$
$$F_{ij} = \sum_p \frac{1}{\sigma_p^2} \frac{\partial f}{\partial a_i} \frac{\partial f}{\partial a_j}$$
$$B_{ijk} = \sum_p \frac{1}{\sigma_p^2} \frac{\partial f}{\partial a_i} \frac{\partial^2 f}{\partial a_j \partial a_k}$$

Refregier, AA, + 2013

# Monte Carlo Control Loops

![](_page_42_Figure_1.jpeg)

Refregier & Amara (2013)

# Ultra Fast Image Generator (UFig)

![](_page_43_Picture_1.jpeg)

Speed the driving factor

As fast as SExtractor (or faster) Subaru Image (0.25 deg2,R~26,10k×8k) generated in: 30sec on a laptop 30µsec per galaxy

HOPE: A Python Just-In-Time compiler for astrophysical computations

Akeret et al 2014 http://hope.phys.ethz.ch

	Python (NumPy)	Numba	Cython	Nuitka (NumPy)	PyPy (NumPy)	numexpr (8 cores)	HOPE	C++
Fibonacci	57.4	65.7 <sup><i>a</i></sup>	1.1	26.7	21.1		1.1	1.0
Quicksort	79.4	b	4.6	61.0	45.8		1.1	1.0
Pi sum	27.2	1.0	1.1	13.0	1.0		1.0	1.0
10 <sup>th</sup> order	2.6	2.2	2.1	1.2	12.1	1.4	1.1	1.0
Simplify	1.4	1.5 <sup><i>ab</i></sup>	1.8	1.4	23.2	0.6	0.015	1.0
Pairwise	1357.8	18	10	1247.7	277.8		17	10
distance	(8.7)	1.0	1.0	(9.5)	(60.4)		1.7	1.0
Star PSF	265.4	250.4 <sup><i>a</i></sup>	46.2	234.6	339.5		2.2	1.0

# Calibrating Shear measurement

![](_page_44_Picture_1.jpeg)

Bruderer et al (2015)

![](_page_45_Figure_0.jpeg)

![](_page_46_Figure_0.jpeg)

![](_page_46_Figure_1.jpeg)

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#### Dark Energy Survey & Latest Results

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Science Verification phase was enormously successful. First results are consistent with expectations and first science survey results are due this year.

Cosmic Concordance?

In the era of precision multiple probes, concordance of the model needs to be tested in a consistent way. Relative entropy is simple and powerful tool for this

#### Precision Measurements Looking Forward

Ambitious hardware driven projects need to be matched by sophisticated analysis methods. In particular systematic use of simulations and forward modelling is crucial.

#### End

Grandis et al (2015)

![](_page_50_Figure_1.jpeg)

![](_page_51_Figure_0.jpeg)