

# The origin of primordial non-gaussianity

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- **How may we compute the non-gaussianity?**

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- Constant value  $\zeta(\mathbf{x})$  at  $T \sim 10$  keV provides the initial condition for adiabatic perturbations.

# The correlators

Spectrum  $\mathcal{P}_\zeta$ , bispectrum<sup>†</sup>  $f_{\text{NL}}$ , trispectrum<sup>††</sup>  $\tau_{\text{NL}}$ :

$$\langle \zeta_{\mathbf{k}} \zeta_{\mathbf{k}'} \rangle = (2\pi)^3 \delta(\mathbf{k} + \mathbf{k}') K_1 \mathcal{P}_\zeta$$

$$\frac{5}{3} \langle \zeta_{\mathbf{k}} \zeta_{\mathbf{k}'} \zeta_{\mathbf{k}''} \rangle = (2\pi)^3 \delta(\mathbf{k} + \mathbf{k}' + \mathbf{k}'') K_2 \mathcal{P}_\zeta^2 f_{\text{NL}}$$

$$\langle \zeta_{\mathbf{k}} \zeta_{\mathbf{k}'} \zeta_{\mathbf{k}''} \zeta_{\mathbf{k}'''} \rangle_{\text{c}} = (2\pi)^3 \delta(\mathbf{k} + \mathbf{k}' + \mathbf{k}'' + \mathbf{k}''') K_3 \mathcal{P}_\zeta^3 \tau_{\text{NL}}$$

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where the kinematic factors depend on the wave-vectors:

$$\begin{aligned}K_1 &\equiv 2\pi^2 / k^3 \\ K_2 &\equiv K_1(k) K_1(k') + 5\text{perms} \\ K_3 &\equiv K_2 K_1(|\mathbf{k} + \mathbf{k}''|) + 23\text{perms}\end{aligned}$$

<sup>†</sup> Komatsu/Spergel 2000; Maldacena 2003    <sup>††</sup> Boubekur/DHL 2005

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- Assume some light fields  $\phi_i(\mathbf{x}, t_1)$  define subsequent expansion  $N(\mathbf{x}, t)$ 
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$$\begin{aligned} \zeta(\mathbf{x}, t) &= N_i(\phi_i(\mathbf{x}), \rho(t)) - N(\phi_i, \rho(t)) \\ &= N_i(t)\delta\phi_i(\mathbf{x}) + \frac{1}{2}N_{ij}(t)\delta\phi_i(\mathbf{x})\delta\phi_j(\mathbf{x}) + \dots \end{aligned}$$

DHL, Malik & Sasaki 05; DHL & Rodriguez 05 (non-perturbative)

# Calculating $f_{\text{NL}}$

Simplest case

$$\zeta(\mathbf{x}) = b\phi(\mathbf{x}) + (\phi(\mathbf{x}) - \bar{\phi})^2$$

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$$= \frac{b^2}{(2\pi)^3} \int \langle \phi_{\mathbf{k}_1} \phi_{\mathbf{k}_2} \phi_{\mathbf{q}} \phi_{\mathbf{k}_3 - \mathbf{q}} \rangle d^3 q + \text{cyclic}$$

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Gives  $\boxed{(3/5) f_{\text{NL}} = 1/b^2}$

# Loop correction

Include quadratic term of  $\zeta$

$$\langle \zeta_{\mathbf{k}} \zeta_{\mathbf{k}'} \rangle = b^2 \langle \phi_{\mathbf{k}} \phi_{\mathbf{k}'} \rangle + \langle [(\phi - \bar{\phi})^2]_{\mathbf{k}} [(\phi - \bar{\phi})^2]_{\mathbf{k}'} \rangle$$

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$$\mathcal{P}_{\zeta} = b^2 \mathcal{P}_{\phi} + 4\mathcal{P}_{\phi}^2 \ln(kL)$$

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Now go to box size  $M \ll L$

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But  $\overline{(\bar{\phi}_M - \bar{\phi})^2} = \mathcal{P}_{\phi} \ln L/M$  hence  $\overline{\mathcal{P}_{\zeta M}} = \mathcal{P}_{\zeta}$



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so  $\partial N / \partial \phi = M_{\text{P}}^{-2} V / V'$  giving

$$\zeta = \left( \frac{\partial N}{\partial \phi} \delta\phi \right) + \left( \epsilon - \frac{1}{2} \eta \right) \left( \frac{\partial N}{\partial \phi} \delta\phi \right)^2$$

where  $\epsilon \equiv (M_{\text{P}}^2/2)(V'/V)^2$  and  $\eta \equiv M_{\text{P}}^2 V''/V$

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How to calculate  $\langle \delta\phi_{\mathbf{k}_1} \delta\phi_{\mathbf{k}_2} \delta\phi_{\mathbf{k}_3} \rangle$

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Keeping just tree level, total  $f_{\text{NL}}$  is [with  $0 < y < 5/6$ ]

$$\frac{3}{5} f_{\text{NL}} = -\frac{1}{4} [n - 1 + (r/8)y(k_1, k_2, k_3)] \sim \pm 10^{-2}$$

Maldacena 2003; Seery/Lidsey 2005

# The curvaton model

Mollerach 90; Linde/Mukhanov 96; DHL/Wands 01; Moroi/Takahashi 2001

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- Curvaton model is not an epicycal!
  - Candidate sRH $\nu$  discovered serendipitously

(Hamaguchi/Murayama/Yanagida 01)

# Prediction of curvaton model

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- If  $\sigma_{\text{os}}$  linear and  $\Omega_\sigma \simeq 1$  then  $f_{\text{NL}} = -\frac{4}{5}$

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## To all

- Prediction generally depends on mean values of light scalar fields in observable Universe

# My main messages

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    - Exception ??: single-component inflation